

## Pure Rotation Spectrum of NNO in the Far Infrared Region

Koichi M. T. Yamada

I. Physikalisches Institut Universität zu Köln  
D-5000 Köln 41, West Germany

Z. Naturforsch. **45a**, 837–838 (1990);  
received April 4, 1990

The pure rotational spectrum of NNO has been observed as an impurity in the NO spectrum which has been recorded with a high resolution Fourier transform spectrometer. The observed high- $J$  transitions in the ground vibrational state were analyzed by a least-squares fit together with the available millimeter and submillimeter wave data. It has been proved that the highly precise data of Maki et al. [3] can be used as a wavenumber standard for the far infrared.

Recently Maki and coworkers have revised the spectroscopic parameters of the NNO molecule intending to supply a good wavenumber standard in the infrared region [1–3]. They have measured the vibration-rotation transitions of this molecule in the infrared region with high precision by using heterodyne technique. They have provided also a very accurate set of the constants for the ground state [3]. However, to the best of our knowledge, the pure rotational transitions in the far infrared (FIR) have not been measured yet. In the course of our measurements of NO [4], we have accidentally found the spectrum of NNO in the region from 20 to 50  $\text{cm}^{-1}$  as an impurity. By analyzing these data together with the available millimeter (mmw) and submillimeter (sub-mmw) wave data [5, 6] up to 552 GHz, we have confirmed the ground state constants of [3].

The measurement has been carried out with a Bruker IFS 120 HR vacuum Michelson spectrometer at the molecular spectroscopy laboratory, Physikalisches-Chemisches Institut, Universität Giessen. The sample of NO was sublimated at a low temperature, in order to prevent contamination by water, and was transferred at a total pressure of about 150 Pa into the 1.8 m long cell sealed with polyethylene windows. The total pressure was around 156 Pa. The spectrum was recorded in the region from 20 to 90  $\text{cm}^{-1}$ , using a myler beamsplitter and InSb detector at liq. He tem-

perature. Small amounts of water, which presumably came out of the cell wall, have been detected. In addition to the NO spectrum very regularly spaced lines were observed, whose separation indicated NNO to be the carrier caused by a small impurity in the NO gas.

Table 1. Measured transitions of NNO in the ground state.

$J' - J''$	Observed	Obs-Calc	Unit
24–23	20.10268	–0.00012	$\text{cm}^{-1a}$
25–24	20.93954	0.00000	$\text{cm}^{-1a}$
26–25	21.77625	0.00006	$\text{cm}^{-1a}$
27–26	22.61283	0.00010	$\text{cm}^{-1a}$
28–27	23.44908	–0.00007	$\text{cm}^{-1a}$
29–28	24.28544	–0.00002	$\text{cm}^{-1a}$
30–29	25.12173	0.00008	$\text{cm}^{-1a}$
31–30	25.95782	0.00012	$\text{cm}^{-1a}$
32–31	26.79374	0.00012	$\text{cm}^{-1a}$
33–32	27.62954	0.00013	$\text{cm}^{-1a}$
34–33	28.46510	0.00003	$\text{cm}^{-1a}$
35–34	29.30064	0.00007	$\text{cm}^{-1a}$
36–35	30.13594	0.00001	$\text{cm}^{-1a}$
37–36	30.97119	0.00005	$\text{cm}^{-1a}$
38–37	31.80625	0.00006	$\text{cm}^{-1a}$
39–38	32.64120	0.00013	$\text{cm}^{-1a}$
40–39	33.47586	0.00006	$\text{cm}^{-1a}$
41–40	34.31028	–0.00008	$\text{cm}^{-1a}$
42–41	35.14466	–0.00007	$\text{cm}^{-1a}$
43–42	35.97887	–0.00008	$\text{cm}^{-1a}$
44–43	36.81301	0.00005	$\text{cm}^{-1a}$
45–44	37.64684	0.00003	$\text{cm}^{-1a}$
46–45	38.48049	0.00004	$\text{cm}^{-1a}$
47–46	39.31394	0.00004	$\text{cm}^{-1a}$
48–47	40.14717	0.00002	$\text{cm}^{-1a}$
49–48	40.98036	0.00016	$\text{cm}^{-1a}$
50–49	41.81312	0.00007	$\text{cm}^{-1a}$
51–50	42.64565	–0.00003	$\text{cm}^{-1a}$
52–51	43.47817	0.00007	$\text{cm}^{-1a}$
53–52	44.31021	–0.00008	$\text{cm}^{-1a}$
55–54	45.97391	–0.00010	$\text{cm}^{-1a}$
57–56	47.63681	0.00002	$\text{cm}^{-1a}$
5–4	125613.69600	–0.00111	$\text{MHz}^b$
6–5	150735.04600	0.00308	$\text{MHz}^b$
7–6	175855.62300	–0.00558	$\text{MHz}^b$
8–7	200975.25600	–0.07138	$\text{MHz}^c$
10–9	251211.55700	–0.00068	$\text{MHz}^b$
12–11	301442.70000	–0.02022	$\text{MHz}^b$
15–14	376777.75300	0.00822	$\text{MHz}^d$
16–15	401885.80200	0.01432	$\text{MHz}^d$
17–16	426991.80800	0.00456	$\text{MHz}^d$
18–17	452095.67000	0.00467	$\text{MHz}^d$
19–18	477197.24700	0.00033	$\text{MHz}^d$
20–19	502296.42300	0.00228	$\text{MHz}^d$
21–20	527393.05100	–0.00977	$\text{MHz}^d$
22–21	552487.03600	–0.00411	$\text{MHz}^d$

<sup>a</sup> Present work: estimated uncertainty of 3 MHz.

<sup>b</sup> From [5]: estimated uncertainty of 10 kHz.

<sup>c</sup> From [5]: estimated uncertainty of 100 kHz.

<sup>d</sup> From [6]: estimated uncertainty of 10 kHz.

Reprint request to Dr. Koichi M. T. Yamada, I. Physikalisches Institut, Universität zu Köln, D-5000 Köln 41, West Germany.

0932-0784 / 90 / 0600-0837 \$ 01.30/0. – Please order a reprint rather than making your own copy.



Dieses Werk wurde im Jahr 2013 vom Verlag Zeitschrift für Naturforschung in Zusammenarbeit mit der Max-Planck-Gesellschaft zur Förderung der Wissenschaften e.V. digitalisiert und unter folgender Lizenz veröffentlicht: Creative Commons Namensnennung-Keine Bearbeitung 3.0 Deutschland Lizenz.

Zum 01.01.2015 ist eine Anpassung der Lizenzbedingungen (Entfall der Creative Commons Lizenzbedingung „Keine Bearbeitung“) beabsichtigt, um eine Nachnutzung auch im Rahmen zukünftiger wissenschaftlicher Nutzungsformen zu ermöglichen.

This work has been digitalized and published in 2013 by Verlag Zeitschrift für Naturforschung in cooperation with the Max Planck Society for the Advancement of Science under a Creative Commons Attribution-NoDerivs 3.0 Germany License.

On 01.01.2015 it is planned to change the License Conditions (the removal of the Creative Commons License condition “no derivative works”). This is to allow reuse in the area of future scientific usage.

	Present work		Maki et al. [3]
	Combined fit	FIR only	
$B$ /MHz	12561.63365 (25)	12561.646 (31)	12561.63395 (10)
$D$ /kHz	5.27882 (51)	5.273 (17)	5.27915 (12)
$H$ /MHz	-0.17 (40)	-1.9 (29)	-4.921 (138)

Table 2. The rotational and centrifugal distortion constants of NNO in the ground vibrational state<sup>a</sup>

<sup>a</sup> Numbers in the parentheses are the standard deviation in units of the last digit quoted.

The spectrum was calibrated using the water lines in the recorded spectrum, the positions of which are listed by Guelachvili and Rao [7], and some sub-mmW lines of NO which were measured by coherent radiation sources [8]. In the recorded region we have identified 77 standard lines, and the fit of those to a non-linear calibration equation,

$$\nu = a \nu_{\text{exp}} + b/\nu_{\text{exp}}, \quad (1)$$

resulted in a standard deviation of 2.3 MHz, which represents the accuracy of the present measurements. The non-linear part in (1) represents the diffraction effect at the aperture of the radiation source which gives a significant effect at long wavelengths [9].

The measured transition wavenumbers of the NNO lines are summarized in Table 1 together with the mmW and sub-mmW data from [5] and [6] which were included in the present analysis. The constants obtained from the weighted least-squares fit are listed in Table 2. We have used the weight reciprocally proportional to the square of the estimated experimental uncertainties: 10 kHz for most of the mmW and sub-mmW lines and 3 MHz for the FIR data. The constants resulting from this fit are listed in Table 2, together with the constants given by Maki and

coworkers [3] and the constants obtained from FIR data only.

The present constants agree within one standard deviation with those of [3]. The constants obtained from the FIR only are also in a good agreement with those obtained from the combined fit. The heterodyne measurements of Maki and coworkers are very precise; the uncertainties for the three constants  $B$ ,  $D$ , and  $H$  are smaller than those of the present work. From the present result we conclude that the transition frequencies calculated from the ground state constants reported in [3] should be used as a wavenumber standard in the FIR region with an accuracy better than  $0.0001 \text{ cm}^{-1}$ .

#### Acknowledgement

Author wishes to express his thanks to Mr. K. Lattner, Universität Giessen, for his excellent assistance in operating the FT spectrometer, and to Dr. G. Winnewisser for his help in the preparation of this manuscript. The work in Köln was supported in part by the Deutsche Forschungsgemeinschaft through SFB-301.

- [1] J. S. Wells, A. Hinz, and A. G. Maki, *J. Mol. Spectrosc.* **114**, 84 (1985).
- [2] M. D. Vanek, M. Schneider, J. S. Wells, and A. G. Maki, *J. Mol. Spectrosc.* **134**, 154 (1989).
- [3] A. G. Maki, J. S. Wells, and M. D. Vanek, *J. Mol. Spectrosc.* **138**, 84 (1989).
- [4] A. Saleck, G. Winnewisser, and K. M. T. Yamada, manuscript in preparation.
- [5] R. Pearson, T. Sullivan, and L. Frenkel, *J. Mol. Spectrosc.* **34**, 440 (1970).
- [6] B. A. Andreev, A. V. Burenin, E. N. Karyakin, A. F. Krupnov, and S. Shapin, *J. Mol. Spectrosc.* **62**, 125 (1976).
- [7] G. Guelachvili and K. Narahari Rao, "Handbook of Infrared Standards", Academic Press, Orlando, Florida, 1986.
- [8] H. M. Pickett, private communication.
- [9] J. W. C. Johns, private communication.